

# Observation of the $\nu_2$ ( $1^- \leftarrow 0^+$ ) inversion mode band in $\text{H}_3\text{O}^+$ by high resolution infrared spectroscopy<sup>a)</sup>

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The potential energy barrier to inversion of the oxonium ion  $\text{H}_3\text{O}^+$  has been a subject of interest for many years.<sup>1-10</sup> We report the first observation of the high resolution gas phase  $\nu_2$  ( $1^- \leftarrow 0^+$ ) band in absorption by diode laser spectroscopy, an initial step in determining the barrier. In Fig. 1 we show the energy levels and potential energy associated with the  $\nu_2$  inversion mode, analogous to the classical case in ammonia. Measurement of the three indicated infrared bands gives the  $0^- \leftarrow 0^+$  inversion splitting from which the barrier can be determined.<sup>11</sup>

Infrared (and NMR) spectra of  $\text{H}_3\text{O}^+$  have been observed for over 30 years<sup>1,12-18</sup> in crystals and solution, and our work supports the earlier  $\nu_2$  assignments in the 1000-1200  $\text{cm}^{-1}$  region. The  $\nu_3$  stretching mode has been observed in the gas phase by Schwarz<sup>19</sup> in low resolution, pulsed radiolysis experiments. Recently, Saykally and co-workers<sup>20</sup> observed this band at high resolution using the ac discharge velocity modulation technique. The importance of  $\text{H}_3\text{O}^+$  in the chemistry and structure of liquids is well established<sup>21,22</sup> especially with regard to hydrogen bonding in biological systems. The  $\text{H}_3\text{O}^+$  ion is also of central importance in the chemistry of ionized gases, e.g., in laboratory discharges and flames,<sup>23,24</sup> in the upper atmosphere,<sup>25</sup> in comets,<sup>26</sup> and in the interstellar medium.<sup>27</sup> The  $\nu_2$  frequencies reported here can be used to observe  $\text{H}_3\text{O}^+$  in these environments, e.g., by infrared astronomy.

The vibration-rotation absorption spectral lines were observed in a 2 kHz ac glow discharge ( $\sim 400$  W) through a 3:1 mixture of  $\text{H}_2/\text{H}_2\text{O}$  gas at 2 Torr that was pumped through a 150 cm water cooled Pyrex cell of 0.9 cm i.d., much as in Ref. 20. The ion mobility of  $\text{H}_3\text{O}^+$  in the gas mixture and the axial electric field ( $\sim 20$ -30 V/cm) were sufficient to observe 1 *f* velocity modulated absorption spectra.

Two laser analytics mesa-stripe geometry tunable diode lasers were used for infrared sources. Three Q branch transitions are shown in Fig. 2 with typical signal to noise ratios and linewidths. The central feature represents a fractional infrared absorption of 0.08%, which is consistent with an ion density of  $\sim 5 \times 10^{10} \text{ cm}^{-3}$ , a 150 cm path length, and an absorption cross section of  $0.1 \text{ \AA}^2$  ( $\mu \sim 0.4 \text{ D}^{28}$  and  $T = 600 \text{ K}$ ). Further details on drift velocity effects in infrared spectra and on velocity modulation spectroscopy are available in Refs. 29 and 30. Three key elements in the success of this experiment were the improved performance of stripe geometry lasers,<sup>31</sup> the sensitivity of velocity modulation techniques,<sup>30</sup> and the accurate theoretical prediction of the spectrum.<sup>32</sup>

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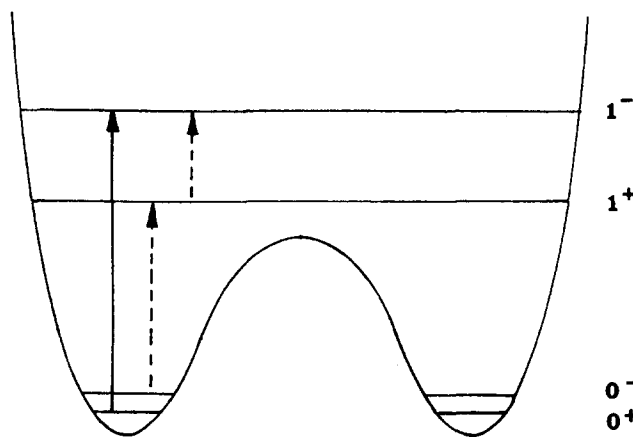


FIG. 1. Sketch of the double well potential for  $\text{H}_3\text{O}^+$  showing the observed band (solid line) and other IR bands (dashed lines).

The observed transition frequencies (with a present, but improvable accuracy of  $\pm 0.005 \text{ cm}^{-1}$ ) are listed in Table I. They fit reasonably well (average deviation =  $0.02 \text{ cm}^{-1}$ ) to the usual parallel band symmetric top expression,<sup>33</sup> with the constants also given in the tabulation. The obviously missing lines are due to gaps in the diodes' frequency coverage. The agreement with the purely *ab initio* calculations<sup>32</sup> of Bunker, Kraemer, and Spirko is striking, but is also consistent with other such successes of high quality *ab initio* calculations on ions.<sup>34-37</sup> The ground state rotational *B* value is in good

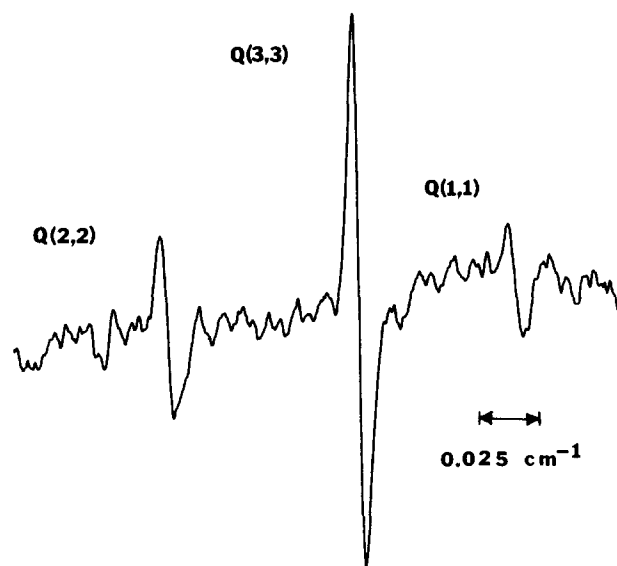


FIG. 2. Q branch transitions of  $\nu_2$  ( $1^- \leftarrow 0^+$ )  $\text{H}_3\text{O}^+$ . Time constant of 1 s, integration of 10 s over full linewidth.

TABLE I. Line positions and molecular constants ( $\text{cm}^{-1}$ ).

$P(5,0)$	831.709	$Q(11,11)$	964.474
$P(5,2)$	834.151	$Q(12,12)$	967.158
$P(4,1)$	858.779	$Q(2,1)$	951.779
$P(4,3)$	864.008	$Q(3,2)$	950.499
$P(3,1)$	884.365	$Q(5,4)$	948.705
$P(3,2)$	886.359	$Q(6,5)$	948.215
$Q(1,1)$	953.975	$Q(7,6)$	947.961
$Q(2,2)$	953.806	$Q(8,7)$	947.981
$Q(3,3)$	953.897	$Q(9,8)$	948.239
$Q(4,4)$	954.254	$Q(10,9)$	948.766
$Q(5,5)$	954.867	$Q(13,12)$	951.764
$Q(6,6)$	955.752	$R(1,0)$	996.064
$Q(7,7)$	956.922	$R(2,2)$	1017.951
$Q(8,8)$	958.354	$R(3,0)$	1033.281
$Q(9,9)$	960.083	$R(3,1)$	1033.963
$Q(10,10)$	962.120	$R(3,3)$	1039.391
	Observed	Predicted <sup>a</sup>	
$\nu_0$	954.417(14)	992	
$B_1$	10.690(4)	10.66	
$B_0$	11.253(4)	11.14	
$(C_1 - B_1) - (C_0 - B_0)$	0.694(15)	0.67	
$D_{1,J}$ <sup>a</sup>	1.8(2)	7.8	
$D_{0,J}$	13(1)	11.0	
$D_{1,JK}$	-1.9(1)	-10.7	
$D_{0,JK}$	-30(4)	-22.5	
$D_{1,K} - D_{0,K}$	-20(10)	-9.2	

<sup>a</sup> Centrifugal distortion constants in units of  $10^{-4} \text{cm}^{-1}$ .

<sup>b</sup> Reference 32.

agreement with the result of  $11.23 \pm 0.11 \text{cm}^{-1}$  from the  $\nu_3$  band spectrum.<sup>20</sup>

Future laboratory experiments will focus upon the other transitions in Fig. 1, as well as further refinement of the present band and the usual isotopic substitution work for a molecular structure. Also, this observation of  $10 \mu\text{m}$  wavelength ion transitions opens up the possibility of subdoppler saturation spectroscopy of ions employing the ultrahigh spectral purity infrared sources which operate in this region.<sup>38</sup>

*Note added in proof:* Just recently Botschwina, Rosnus, and Reinsch<sup>39</sup> have theoretically predicted the present  $\mu_2$  band origin to be at  $961 \text{cm}^{-1}$ , which is in excellent agreement with our results.

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